#### 10.3.1.1 Excitation and radiation of spectra

### 10.3.1.1.1 Plasmas

A *plasma* of the type occurring in spectrochemical radiation sources may be described as a gas which is at least partly ionized and contains particles of various type, viz. electrons, atoms, ions and molecules. The plasma as a whole is electrically neutral.

A plasma which is in *thermodynamic equilibrium* can be characterized by a single temperature, called the *thermodynamic temperature*. This temperature describes the energy distribution of all particles, the state of ionization, the abundance of the chemical species and the spectral energy distribution through *Maxwell-Boltzmann's law*, *Saha-Eggert's law*, the *law of Mass Action* and *Planck's law* respectively. In practice, serious departures from thermodynamic equilibrium may exist, and different temperature values will be found according to the types of measurement. Several types of "temperature" may then be used for describing the state of plasma:

*Radiance temperature* is the temperature of a black-body radiator that has the same spectral radiance, according to Planck's law, as the radiator considered.

*Electron temperature* is the temperature that describes, through Maxwell's law, the kinetic energy distribution of the free electrons.

*Gas temperature* describes in a similar way the kinetic energy distribution of the gas atoms.

*Excitation temperature* is the temperature that describes, through Boltzmann's law, the relative population distribution of atoms or molecules over their energy levels. We distinguish between *electronic excitation temperature*, *vibrational temperature*, and *rotational temperature*.

*Ionization temperature* is related to the ionization equilibrium described by the Saha-Eggert equation.

With a plasma in thermodynamic equilibrium, all these temperatures should be equal. When only the radiance temperature deviates from the others, we speak of *thermal equilibrium*. In this case, all degrees of freedom of all particles are in equilibrium with each other, but not with the radiation field. The particles then emit *thermal radiation*, but not black-body radiation. When the temperature that describes such a state of thermal equilibrium varies with position inside the source, we speak of *local thermal equilibrium* (LTE), characterized by a *local (gas) temperature*.

*Norm temperature* is the temperature of a plasma at LTE for which a spectral line has maximum spectral radiance. This is the result of the competing effects of ionization and excitation processes.

Туре	Sort	Gas Pressure Range	Types of Spectra Radiated*
Arcs	dc	atmospheric (~100 kPa)	a,m,c
	ac	atmospheric	a,m,c
	current-carrying plasmas	atmospheric	a,m,c
	non-current carrying plasmas	atmospheric	a,m,c
	high pressure	100-600 kPa	c, broad a
	low pressure	10-100 kPa	a,m,c
Lasers	continuous	-	a,m (primary) a,m,c (secondary)
	pulsed	-	a (primary)
Low pressure electrical discharge	arc lamps	1-100 kPa	a,m,c (secondary)
	Geissler	1-10 kPa	a,c
	glow discharge; hollow cathode	0.1-10 kPa	a,m,c
	plane cathode	0.1-10 kPa	a,m,c
Microwave plasmas	loaded line	atmospheric	a,m,c
	induced (electrodeless)	low or atmospheric	a,m,c
Radiofrequency plasmas	inductively-coupled	atmospheric	a,m,c
	capacitively-coupled	atmospheric	a,m,c
Sparks	high voltage	atmospheric	a,m,c
	medium voltage	atmospheric	a,m,c
	low voltage	atmospheric	a,m,c
	vacuum sparks	<10 Pa	a,m,c

 TABLE 10.5
 General classification of radiation sources

\*a - atomic (neutral or ionized) m - molecular (neutral or ionized)

c - continuous

Term	Symb	ol	Practical Unit*
analytical calibration function		x=g(c)	same as x
	or	x = g(q)	same as x
analytical evaluation function		c=f(x) q=f(x)	same as c same as q
atomic number	Ζ	/	1
atomic mass unit	т		1
atomic mass (of nuclide B <sub>x</sub> )	$m_{\rm a}(B_{\rm X}$	<u>(</u> )	g
atomic number of species X	Z(X),Z	$Z_{\rm X}$	1
Avogadro constant	$N_{\rm A}$		mol <sup>-1</sup>
Boltzmann constant	k		J K <sup>-1</sup>
capacitance	С		F
charge density	ρ		C m <sup>-3</sup>
conductance (electrical)	G		$S \cong \Omega^{-1}$
current density	j		$A m^{-2}$
electric current	Ι		А
electric field strength	Ε		$V m^{-1}$
electron current density	J		$A m^{-2}$
electron mass	m <sub>e</sub>		g
electron temperature	T <sub>e</sub>		Κ
elementary charge	е		С
energy	Ε		J
excitation energy	$E_{\rm exc}$		J, ev
excitation energy of state q of species X	$(E_q)_X$		J, ev
excitation potential	$V_{\rm exc}$		V
excitation temperature	$T_{\rm exc}$		K
flow rate	ρ		Pa m <sup>3</sup> s <sup>-1</sup>
frequency (electrical)	f		Hz
frequency (in optical spectroscopy)	ν		Hz
frequency of spectral line emitted due to transition $u \rightarrow 1$	$V_{\rm ul}$		Hz
gas constant	R		J K <sup>-1</sup> mol <sup>-1</sup>
gas temperature	$T_{g}$		K

TABLE 10.6 Terms, symbols and units for measurable quantities for radiation sources

Term	Symbol	Practical Unit*
inductance	L	Н
ionization energy	$E_{\rm ion}, E_{\rm i}$	J, eV
ionization potential	$V_{\rm ion,i}$	V
ionization temperature	$T_{\rm ion}$	K
irradiance	Ε	$W \text{ cm}^{-2}$
kinetic energy of particle	$E_{\rm kin}$	J, eV
magnetic flux	$\phi$	Wb
magnetic flux density	В	Т
mass	m	kg
metastable excitation potential	$V_{\rm meta}$	V
number of particles	N	1
number of particles of species X	$N(X), N_X$	1
number density: (number of particles per unit volume)	п	cm <sup>-3</sup>
number density of particles in state q	$n_{\mathrm{q}}$	cm <sup>-3</sup>
number density of element as free atom	n <sub>at</sub>	cm <sup>-3</sup>
number density of element as free ion	n <sub>ion</sub>	cm <sup>-3</sup>
number density of electrons	n <sub>e</sub>	cm <sup>-3</sup>
number density, <i>total</i> , of element in different forms (atom, ion, molecule) in the gaseous state	<i>n</i> t	cm <sup>-3</sup>
number density of ground state species	X <sub>o</sub>	cm <sup>-3</sup>
number density: n <sub>at</sub> for atoms of X	$X, n_{\rm at}(X)$	cm <sup>-3</sup>
number density: $n_{ion}$ for ions of $X^+$	$X^+, n_{\mathrm{ion}}(X)$	cm <sup>-3</sup>
number density: nt for element X	$X_{t}, n_{t}(X)$	cm <sup>-3</sup>
number density of excited species X*	X*	cm <sup>-3</sup>
number of oscillations	Р	1
oscillator strength for absorptions by transitions from states $1 \rightarrow U$ (lower $\rightarrow$ upper)	$f_{ m IU}$	1
partial pressure	Р	Pa
partial pressure of species X	$p(X), p_X$	Pa
partition function	Z	1
partition function of species X	$Z(X), Z_{\rm X}$	1
period	τ	S

Term	Symbol	Practical Unit*
Planck constant	h	Js
potential difference	U	V
power	Р	W
quantity of electricity	Q	С
radiant energy density	и	J cm <sup>-3</sup>
radiant intensity	Ie	W sr <sup>-1</sup>
radiant power	$\phi$	W
reactance	X	Ω
re-ignition potential	$U_{ m z}$	V
relative intensity of spectral line emitted by transitions from state $u \rightarrow 1$	$I_{ m ul}$	1
spark duration	θ	S
spark gap potential	$U_{g}$	V
spectral radiance	$L_{\lambda}$	$W \text{ cm}^{-2} \text{ nm}^{-1} \text{ sr}^{-1}$
sputtering rate	$q_{ m sp}$	g s <sup>-1</sup>
sputtering yield (number of atoms sputtered per incident ion)	$S_{ m sp}$	1
statistical weight of state q	$g_{ m q}$	1
statistical weight of ground state	$g_{0}$	1
statistical weight of state q of species X	$g_q(X),(g_q)_X$	1
thermodynamic (absolute) temperature	Т	Κ
time	t	S
total gas pressure	$p,p_t$	Pa
transition probability for absorption (by transition from 1 to u state)	$B_{lu}$	$s^{-1} J^{-1} cm^3 nm$
transition probability for spontaneous photon emission (by transition from u to 1 state)	$A_{ m ul}$	s <sup>-1</sup>
wavelength	λ	nm
wavenumber	$\widetilde{v}$ , $\delta^{\!**}$	cm <sup>-1</sup>
work function	$\phi$	V

\* Practical units are decimal multiples or fractions of SI units

\*\*  $\widetilde{v}$  is only used for  $v/c = 1/\lambda$  in vacuum

### 10.3.1.1.2 Pressure effects

The pressure within a plasma influences the characteristics of the radiation, in that it affects the *mean free path* of the particles and their *collisional cross-sections* and the number of collisions per unit of time.

Sources may broadly be classified into 4 different groups, depending on the average gas pressure at which they operate, viz.

Low pressure (<10 kPa) Medium pressure (10-100 kPa) Atmospheric pressure (~100 kPa) High pressure (>100 kPa)

# 10.3.1.1.3 Collisional processes

A particle (atom or molecule) characterized by its collisional cross-section can undergo a change in its state of excitation as a result of collisional processes with other particles. In an *elastic collision*, an exchange only of kinetic energy takes place between the colliding species; in an *inelastic collision* there is an interchange between the kinetic energy and internal energy of the particle.

In *quenching*, a particle in an excited state may lose its energy by *collisional de*excitation.

*Energy transfer* from a particle in a higher state to one in a lower state may occur. A *metastable state* (level) is any excited state which, in principle, by virtue of the selection rules, cannot radiatively combine with any lower state. These metastable states usually have considerably longer lifetimes than ordinary excited states.

## 10.3.1.1.4 Radiative processes

The change in the internal energy of a particle may also be due to *radiative processes*, i.e., the emission or absorption of a photon. A particle in an excited state may undergo a transition to a lower energy level by emission of a photon. This is known as *radiative de-excitation*. If such a transition occurs spontaneously, its probability per second for a given excited particle is termed the *transition probability for spontaneous emission*.

A particle in the *ground state* or an excited state may undergo a transition to a higher energy level by *absorption* of a photon. This is known as *radiative excitation*. For a given particle in the lower state, the probability per second of such a transition in a field with a continuous spectrum is proportional to the *spectral radiant energy density* of the field at the wavelength of the absorption line. The proportionality constant is termed the *transition probability for absorption*. The *transition probability for stimulated emission* is defined in a similar way for the reverse radiative de-excitation process that is induced by the same radiation field.

The *oscillator strength* for absorption is often used instead of the transition probability for the (reverse) spontaneous emission process to which it is proportional. (Note: oscillator strength is the historical term relating to the Rutherford atomic model.)

*Atomic fluorescence* is a combined process of photon absorption by an atom followed by spontaneous photon emission.